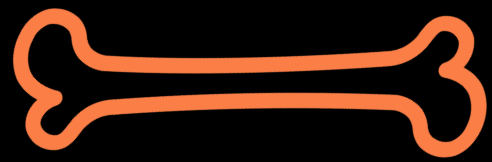
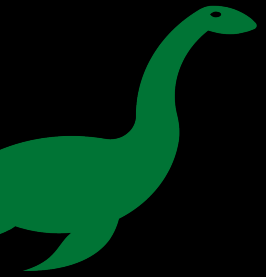


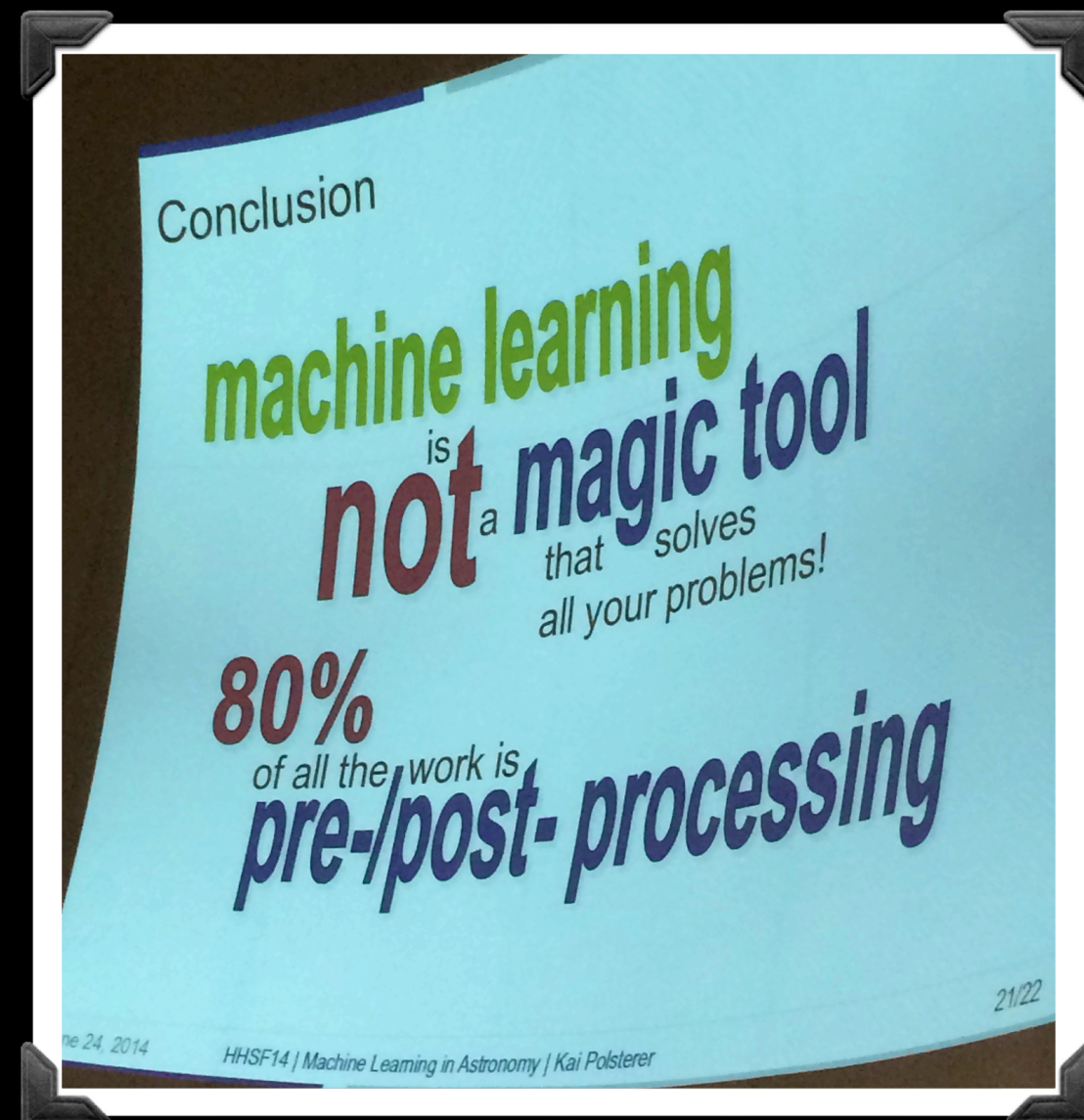
A decade of HHSF 2014-2024, and what's next?

Alyssa Goodman, Harvard

2014



we ♥ filaments



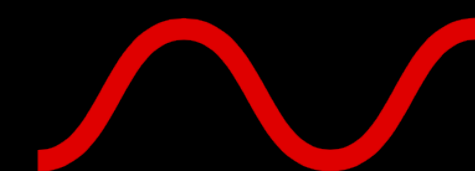
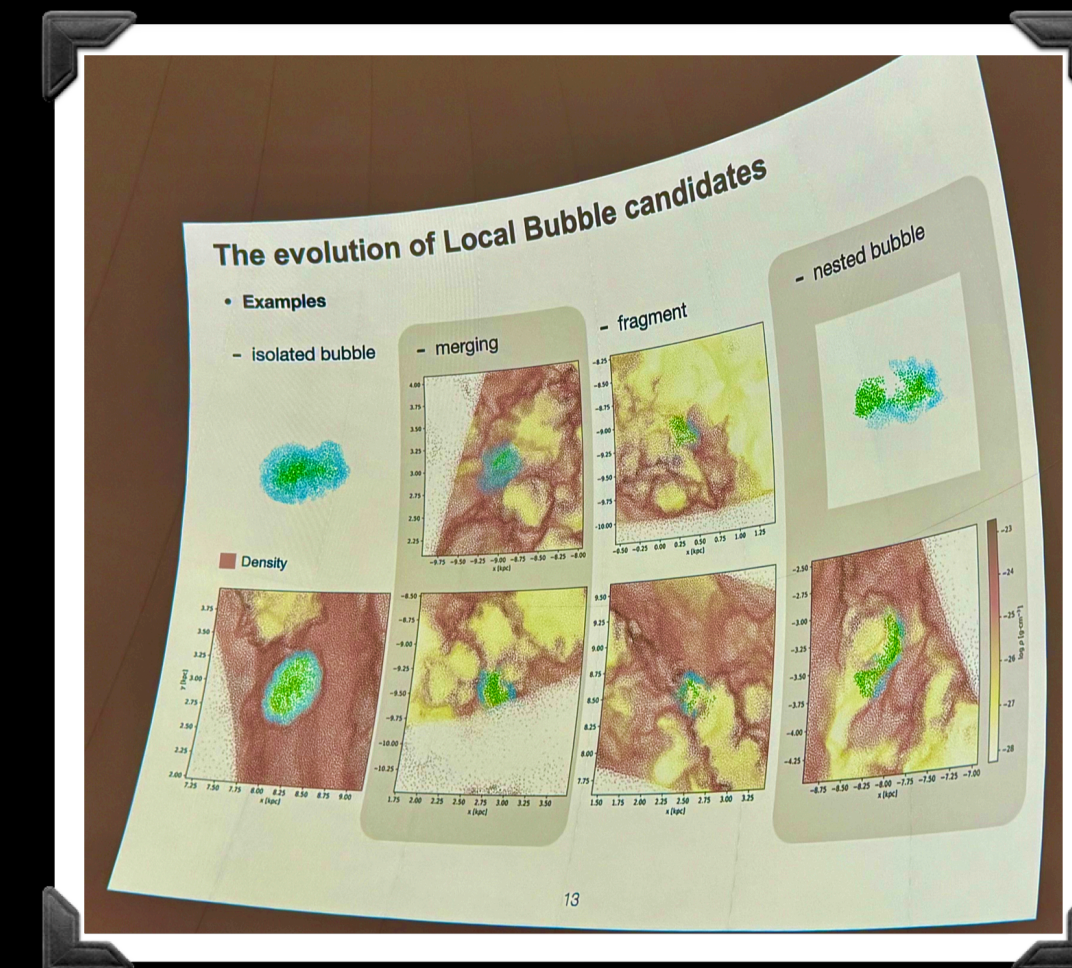
2020



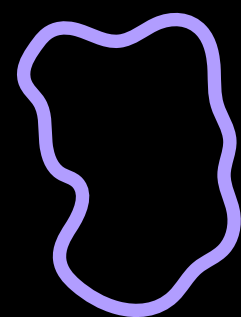
2022



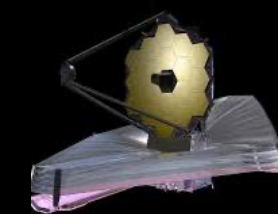
2024



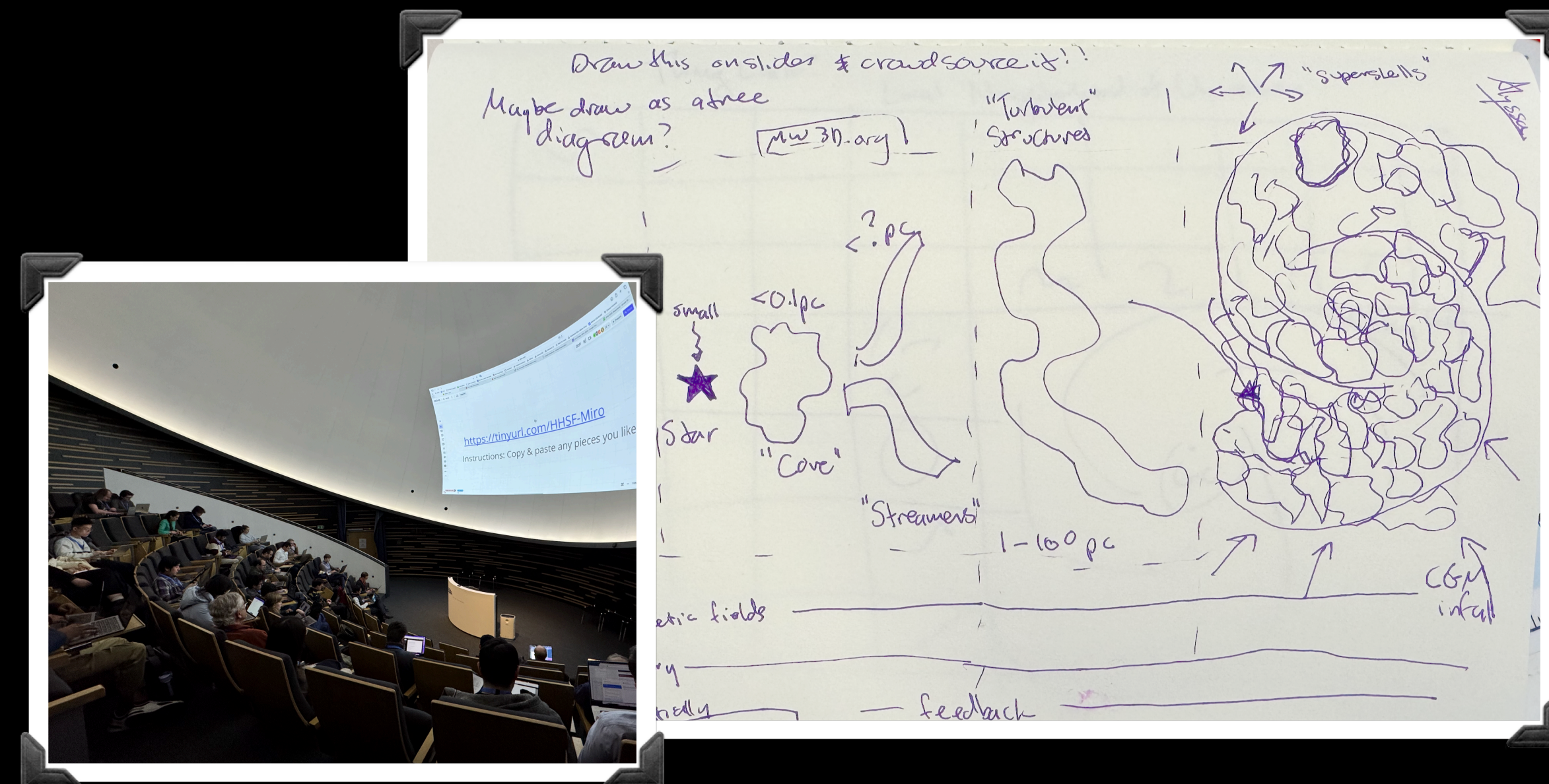
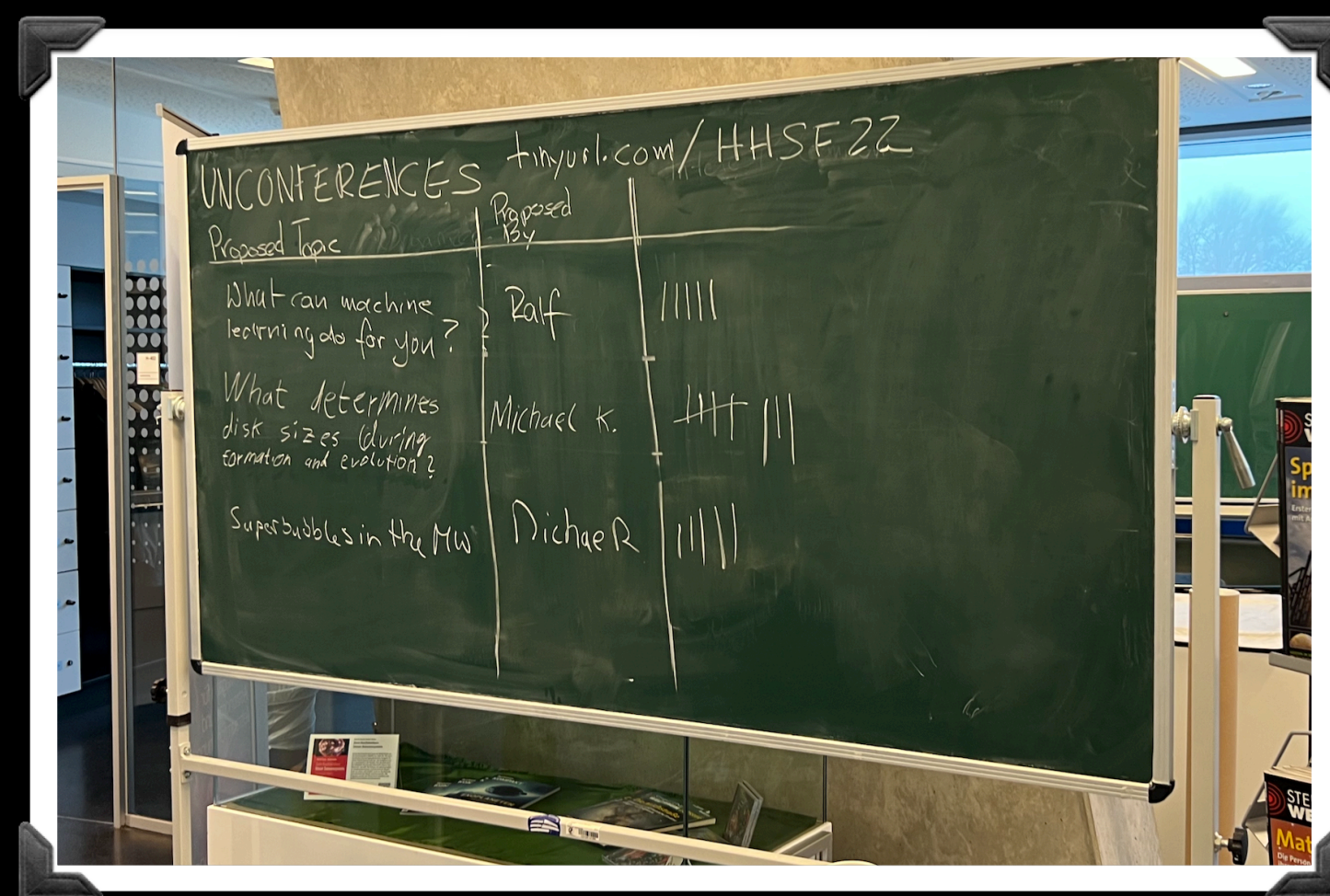
connecting to larger scales



we ♥ bubbles & JWST!

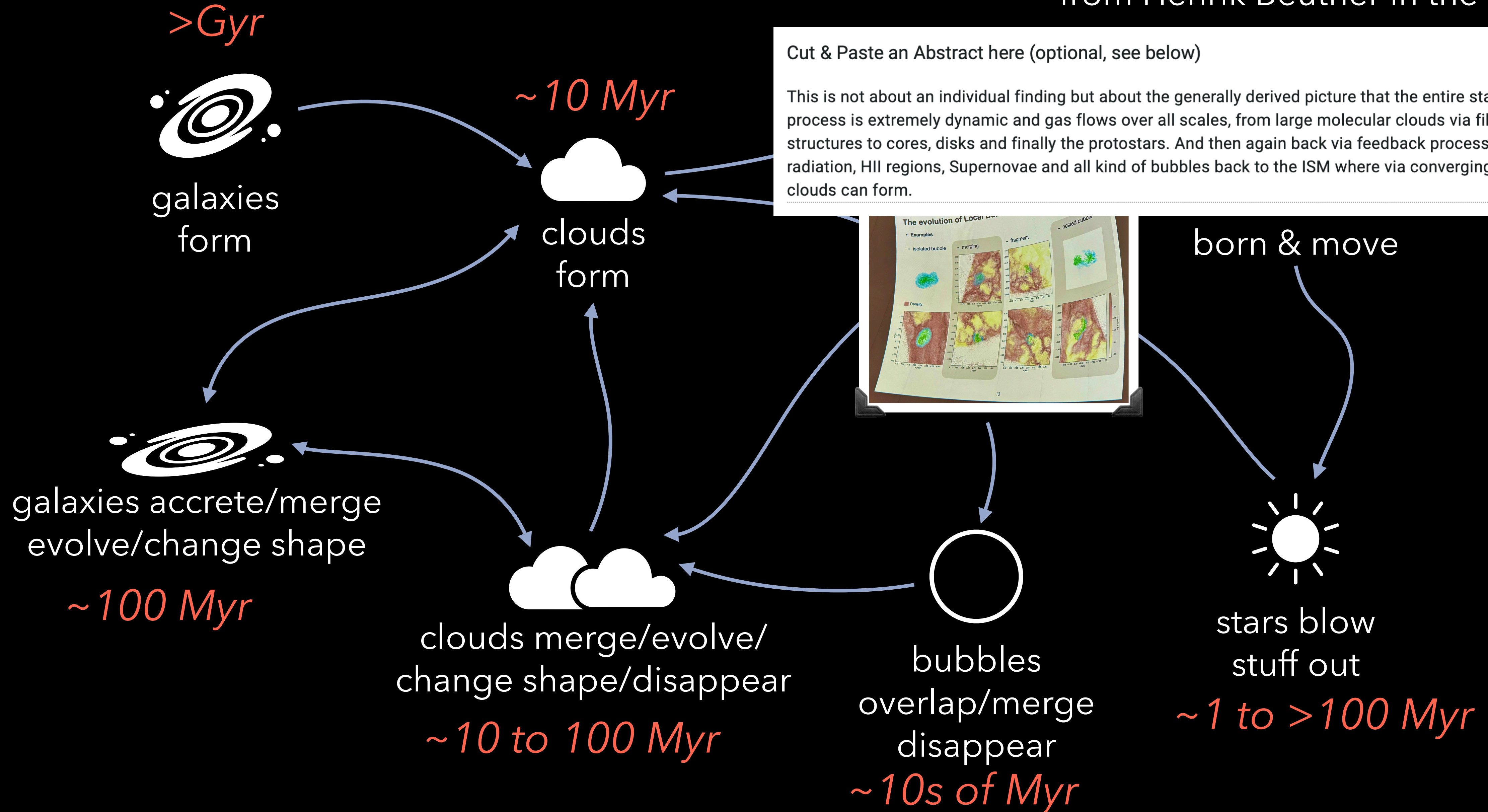


"Everything everywhere all at once?..."



“Everything, Everywhere, All at Once”

from Henrik Beuther in the survey...



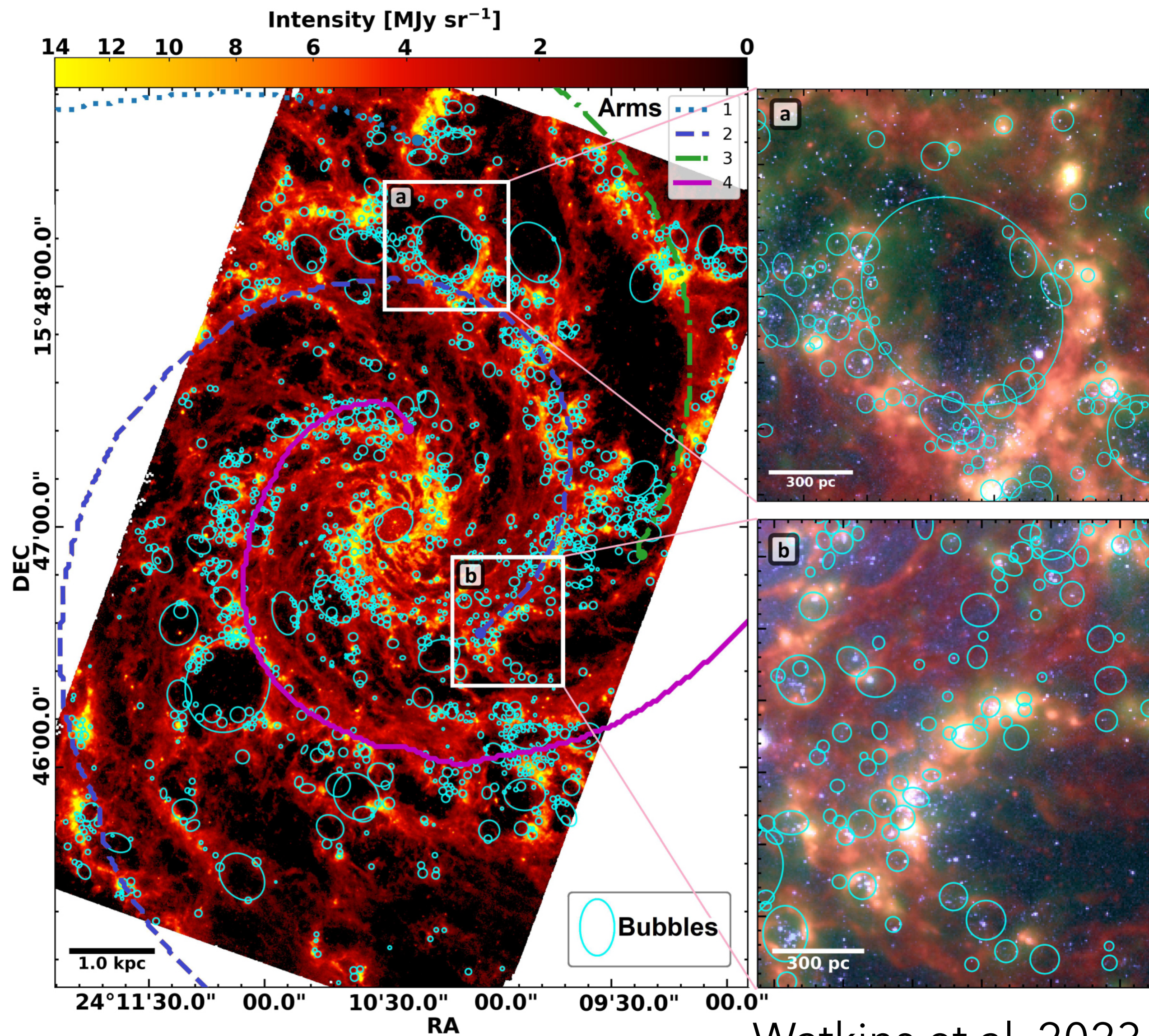
“Everything, Everywhere, All at Once”

1. molecular clouds *“form”*
2. new stars *form* in “molecular clouds”
3. stars *move* while and after they form
4. clouds *“evolve”* (grow, shrink, merge, change shape, change density) over time
5. winds & explosions from stars (called “feedback”) *cause* some of the clouds’ evolution
6. the most “significant” feedback events can *cause* new molecular clouds to form (see #1!)



Nothing stays still. Nothing lasts forever.

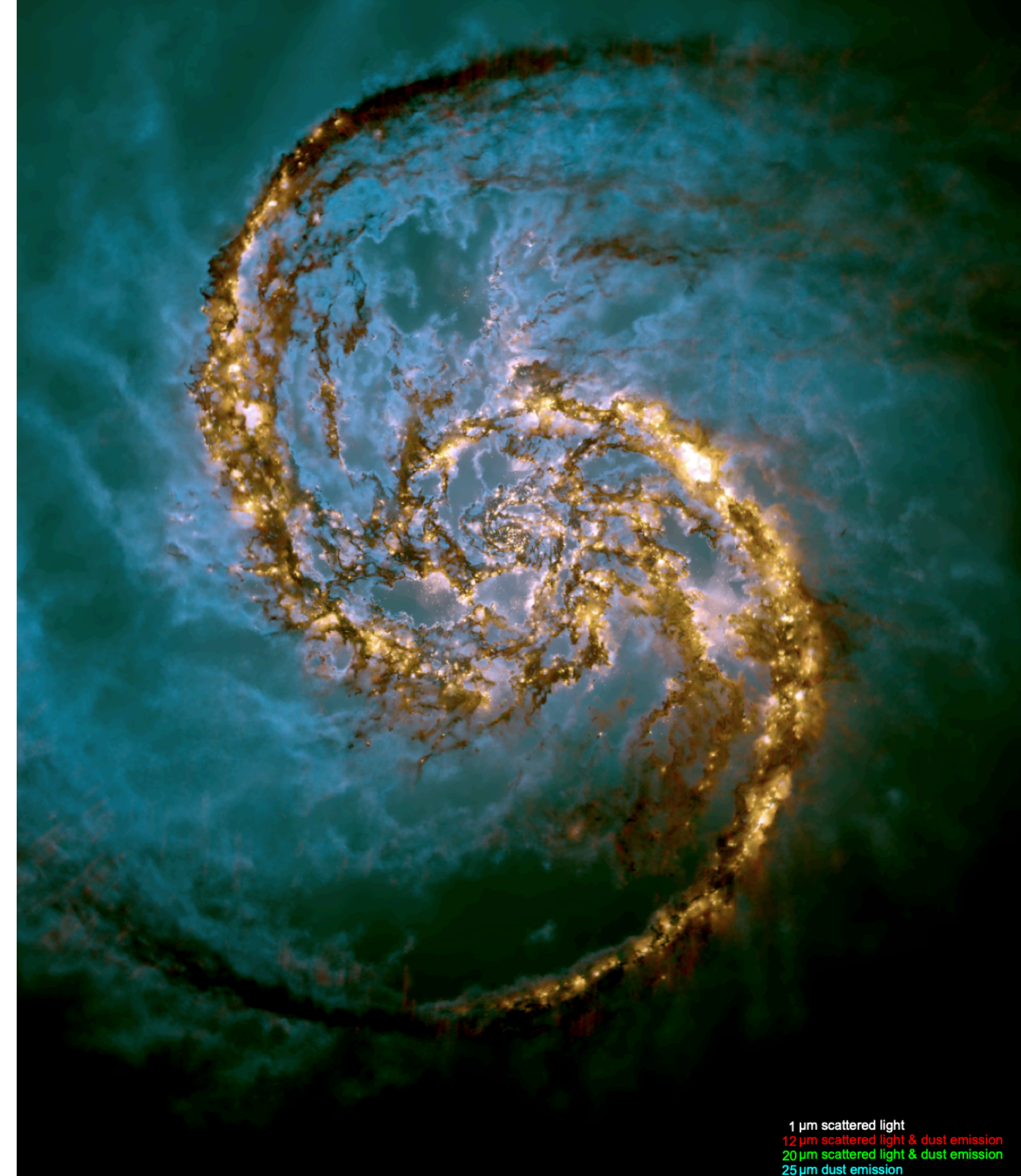
NGC 628 from JWST



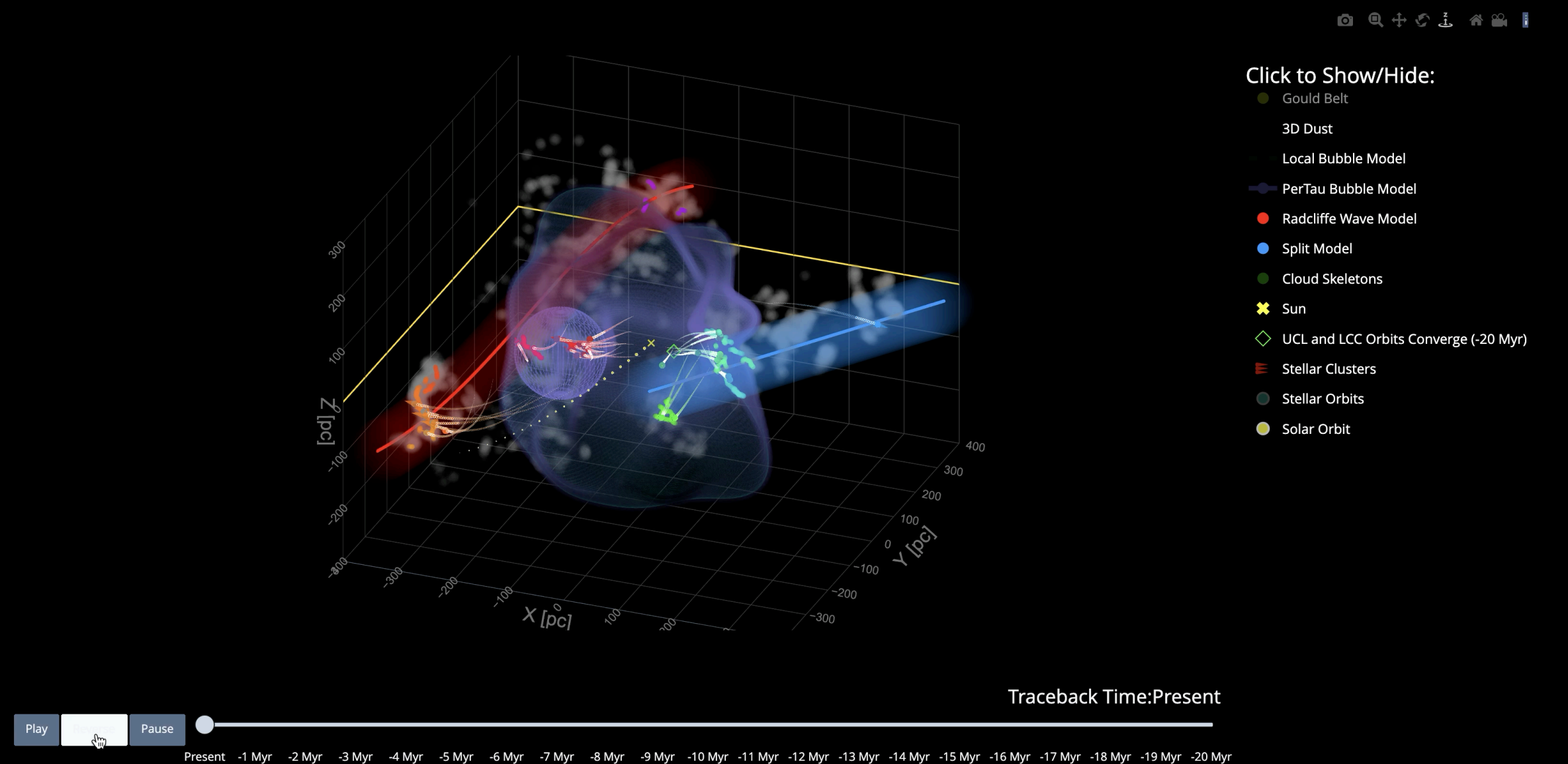
Watkins et al. 2023

The Whirlpool-Galaxy M51 modelled on a Computer

Stefan Reißl & Robin Treß

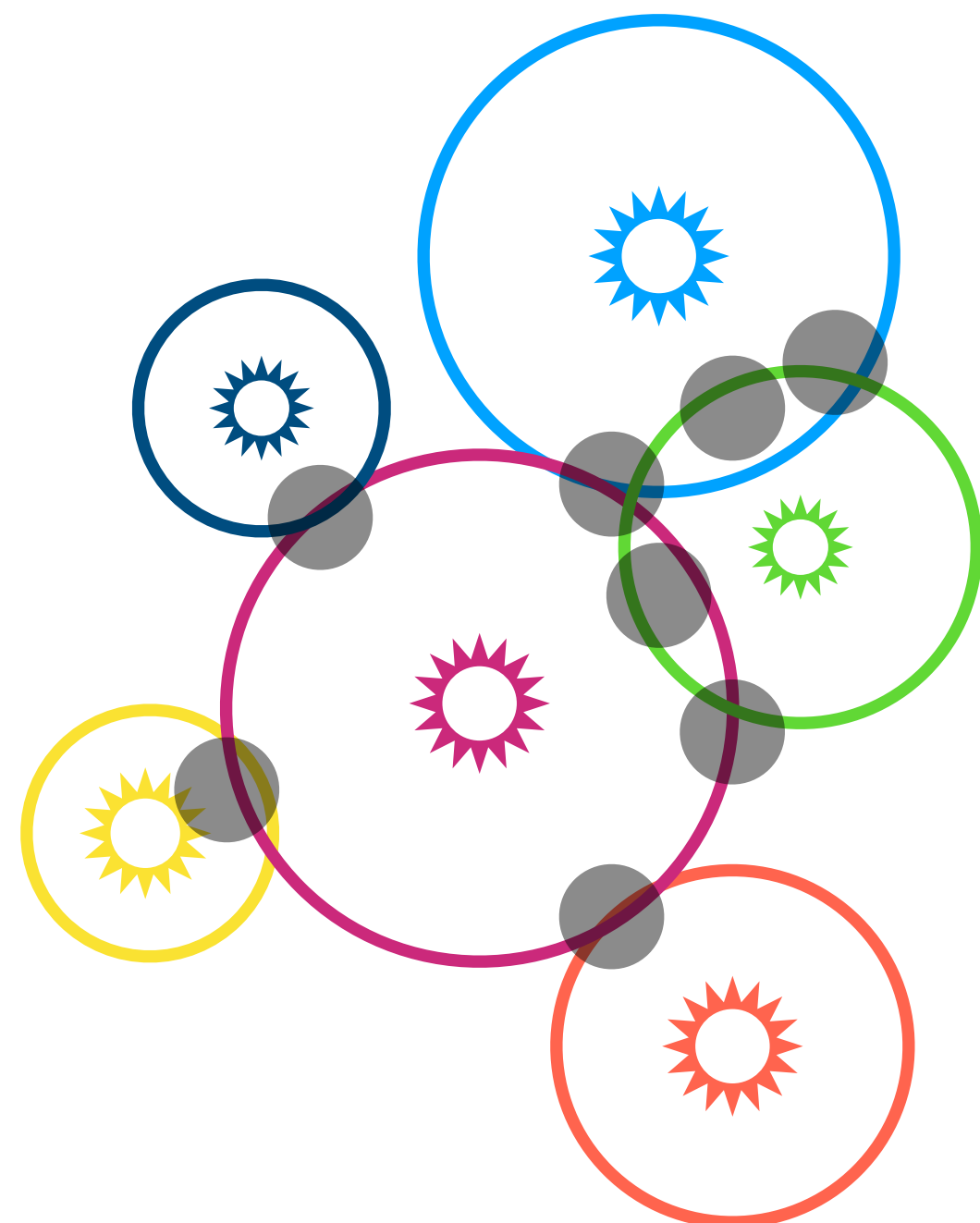


Do filament "collisions" cause bubbles?



https://faun.rc.fas.harvard.edu/czucker/Interactive_Figure3_Alyssa.html,
based on Zucker et al. 2022

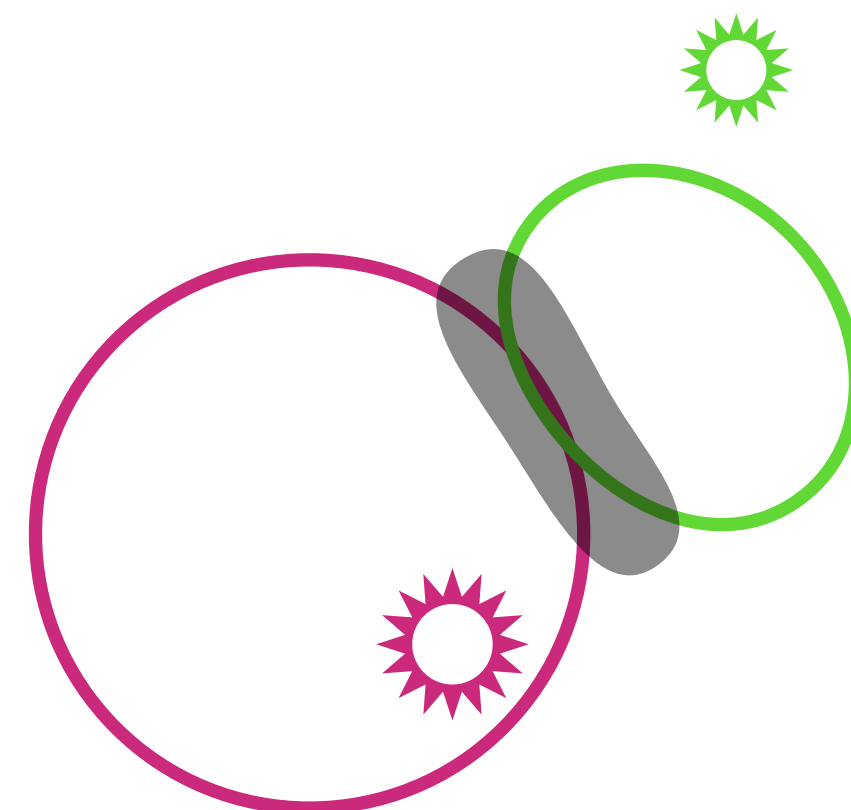
Thought Experiments on the Relationship between Feedback and Dense Star-Forming Gas



Panel 1: Idealized situation (unrealistic).

All cavities are circles. Progenitor star or cluster centered and still present, and all feedback events were roughly simultaneous, and of only slightly different energy. "Spherical" dense clouds form at bubble overlaps.

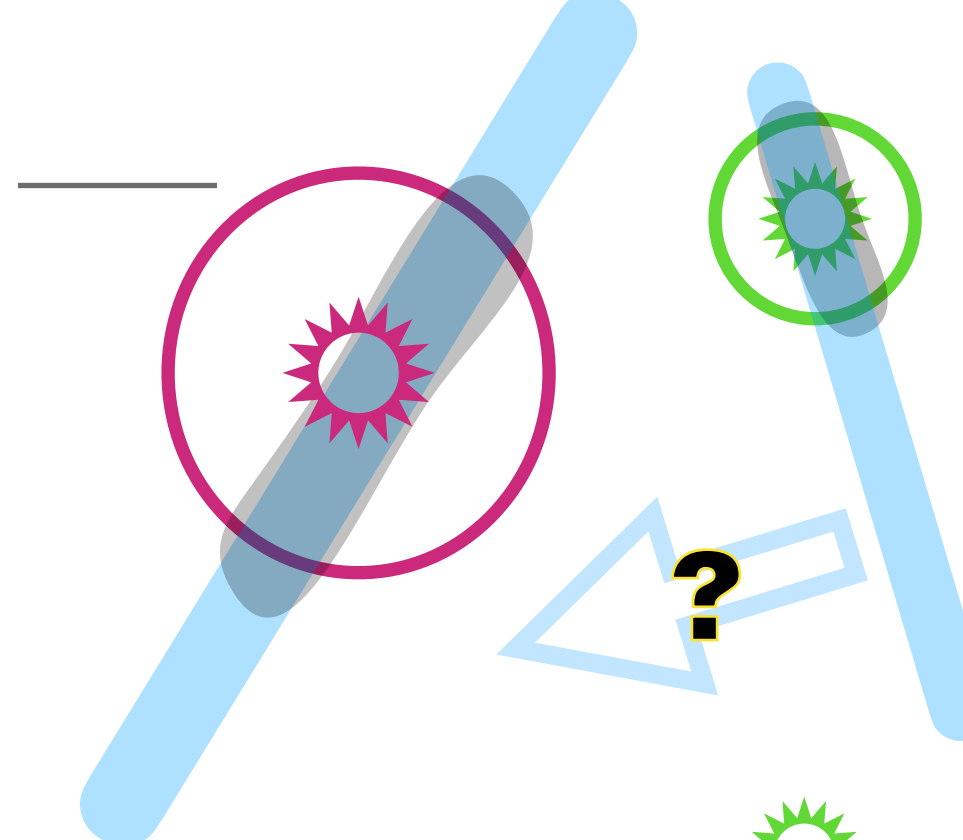
Notes: This is a 2-D cartoon of a 3D process. In all panels, circles or ellipses represent *expanding* cavities of low-density, caused by feedback events, typically explosions of one or more supernovae, originating in a stellar cluster represented by a ☀ symbol. Semi-transparent shapes are meant to represent highly-idealized "molecular clouds" or pieces of larger-scale Galactic structures.



Panel 2: Slightly less idealized situation, shown for just two cavities, for simplicity.

Cavities distorted by asymmetries, especially in ambient medium, some progenitor stars or clusters can still be identified, and the two feedback events are assumed have overlapped in time enough to form dense, non-spherical, clouds at bubble overlaps.

"Past"



Present

Panel 3: Imagining the past and dense clouds' origins.

When one considers the origins of feedback progenitors' dense-cloud homes, things get complicated, due to the limited lifetimes of large star-forming clouds, and, in more realistic scenarios than drawn, mixing of structures with each other over time.

tinyurl.com/HHSF24-survey-results

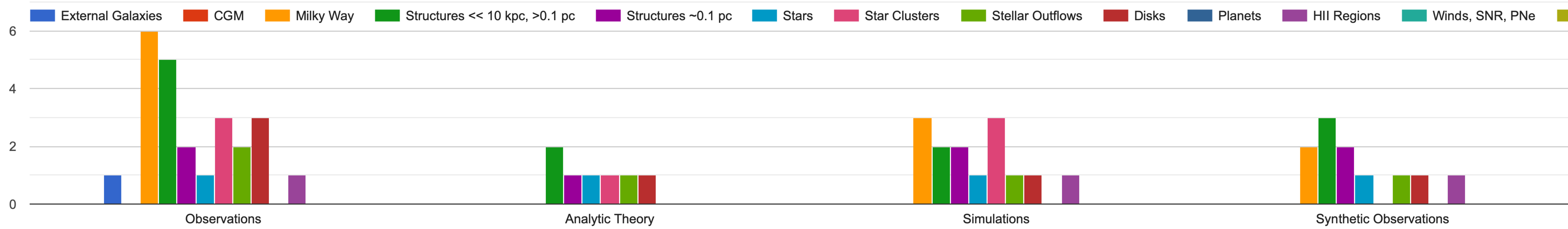
HHSF24 Participants' Favorite Findings and Codes, 2014-2024

Use this form at least **once**, and up to five times, to let everyone know about your favorite **finding** or useful **code** from the time period 2014-2024.

- A "**finding**" can be based on observations, simulations, or statistical or analytic calculations.
- A "**code**" refers to any kind of new or improved software, web technology, or tool, e.g. for simulation, segmentation, data analysis, data visualization, data hosting, AI, etc.

It is fine to submit work in which you were involved—but that is not required, especially if you make multiple entries.

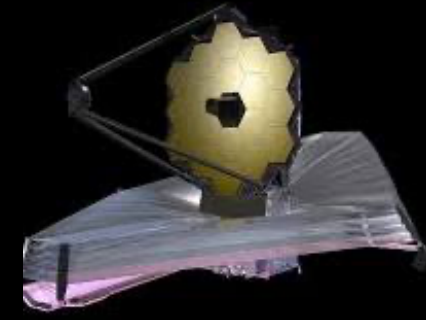
Part 1: Theory, Observation, Stats



PNe Other



JWST image of Fomalhaut [Gaspar 2023 @25.5um] -- multi-eccentric ring disk. Relic of protoplanetary disk or planet-driven asymmetries?



JWST image of HH211

Filament/superbubble structure in surveys, JWST observations, and in galactic multiscale simulations

bubbles, bubbles in JWST images of nearby galaxies

Most nearby young star clusters formed in three massive complexes
Discovery of the Radcliffe Wave cluster census with Gaia
bubbles, the Local Bubble

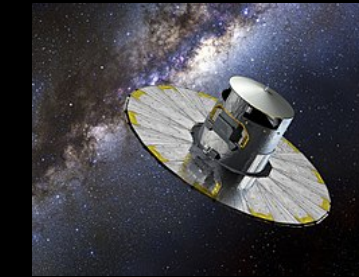


Image matching with astroHOG

Cosmic rays in Arepo
glue

POLARIS

Torch framework for hybrid MHD and N-body simulations

FilFinder

SMA & Pan-STARRS image of IRAS23077+6707

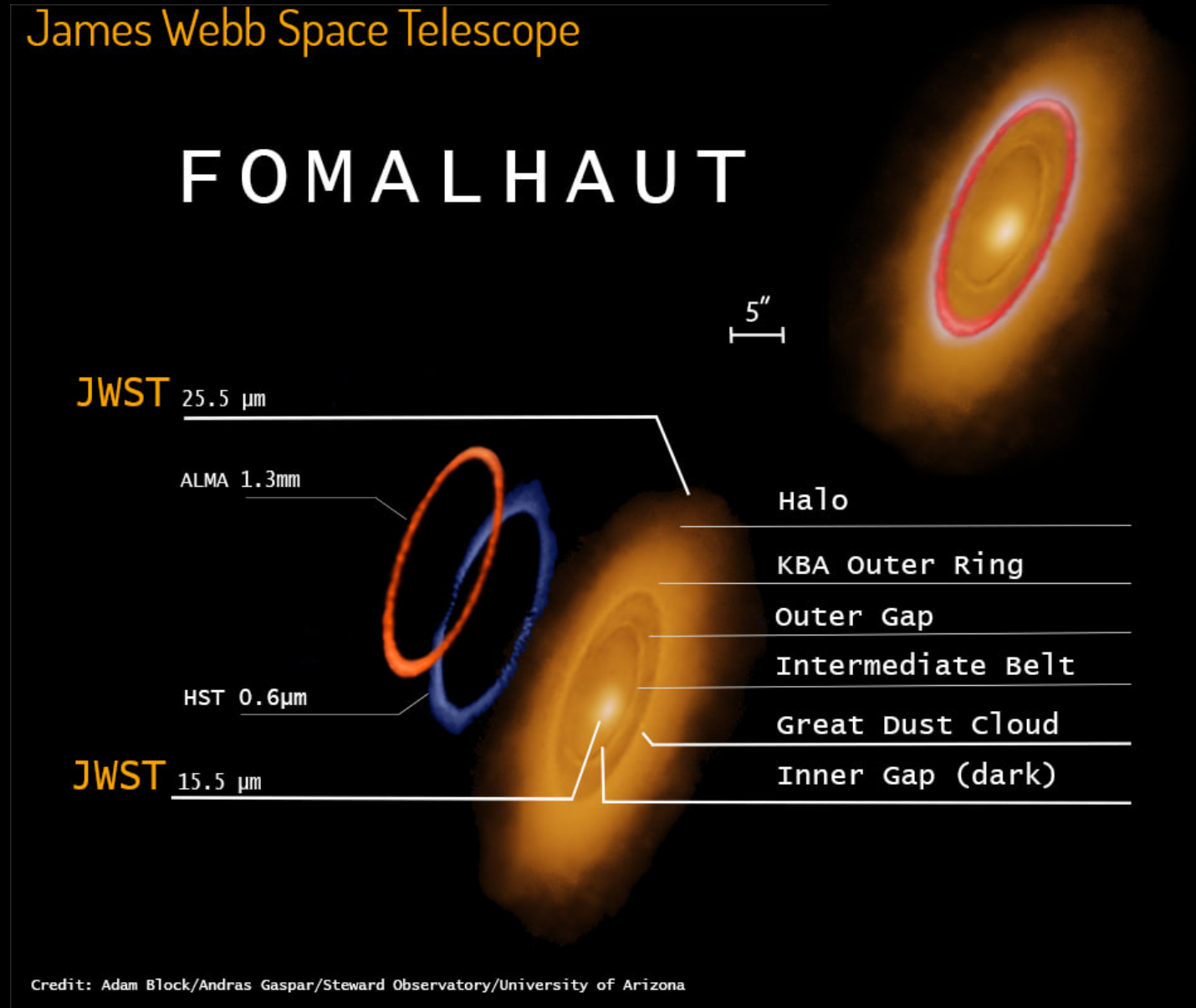
Everything flows

The archaeological record of the Milky Way written in neutral atomic hydrogen (HI)

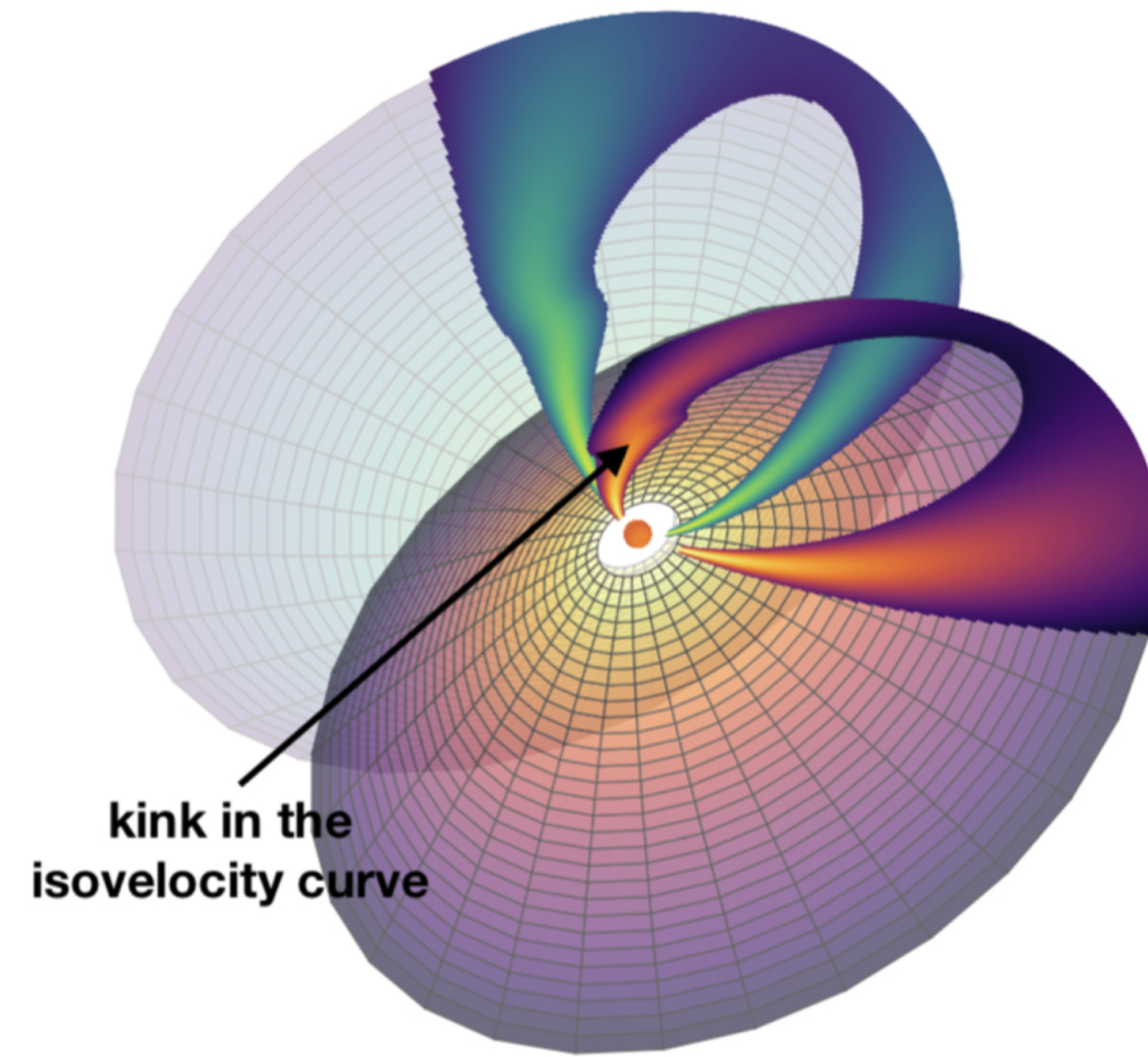


ALMA image of star-forming regions
Kinematic Planetary Signatures (KPS or Kinks)

JWST image of Fomalhaut -- multi-eccentric ring disk.
Relic of protoplanetary disk or planet-driven asymmetries?



Kinematic Planetary Signatures (KPS or Kinks)



SMA & Pan-STARRS image of IRAS23077+6707

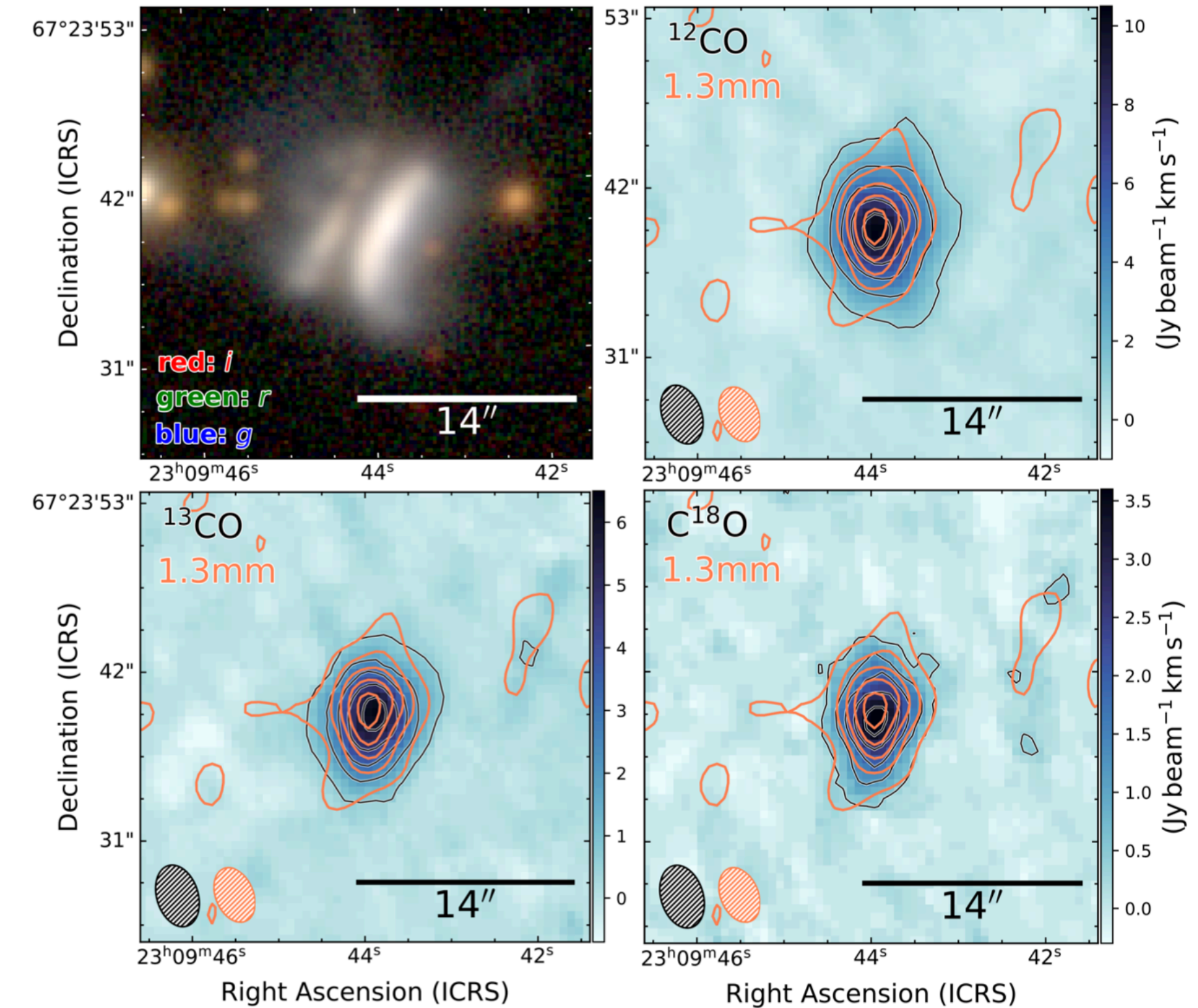


Figure 1. Comparison of optical scattered light and (sub-)mm emission of IRAS 23077+6707. *Top left:* PS1 *irg*-composi image, showing the optical scattered light emission at $\lambda_{\text{mean}} = 7545, 6215$ and 4866 \AA , respectively. *Top right:* SMA ^{12}C integrated intensity (or ‘moment 0’) map, overlaid with its respective contours, as well as the 1.3 mm continuum contours, draw at the 10%, 30%, 50%, 70% and 90% levels of their corresponding maximum emission. *Bottom left:* Same for ^{13}CO . *Bottom right:* Same for C^{18}O . In the lower right of each panel, we show 14" scale bars, and the effective CO (black) and continuum (orange) beams in the lower left of the SMA image.

cf. Gaspar 2023 @25.5 μm

Pinte et al 2023

Monsch et al. 2024

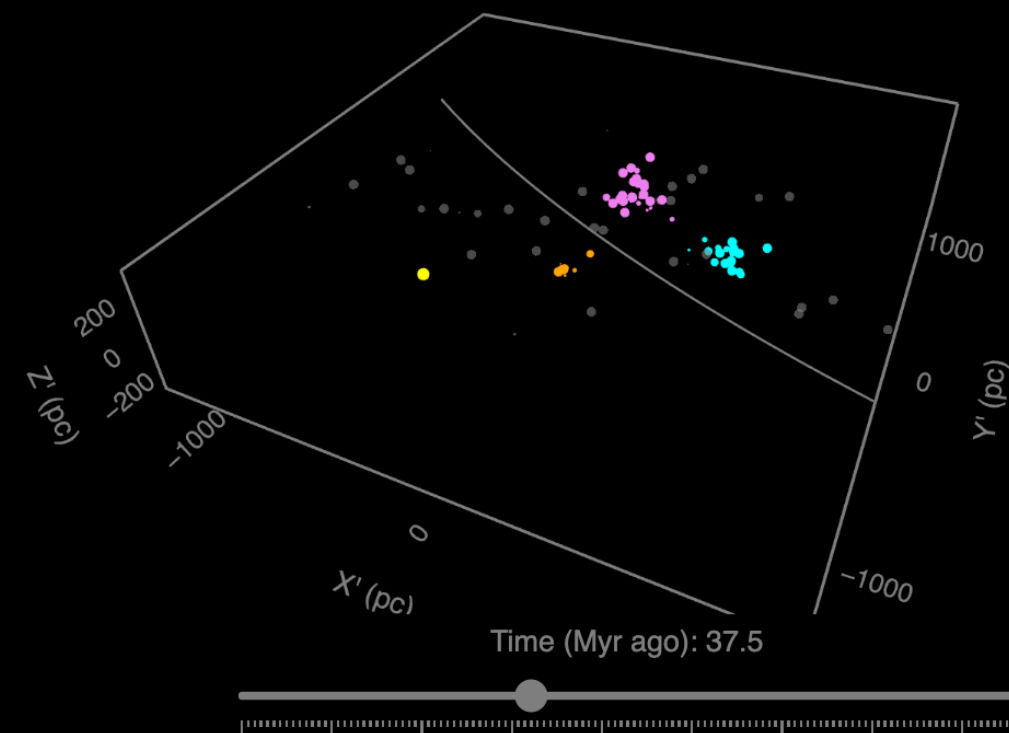
Most nearby young star clusters formed in three massive complexes

Trace young star clusters backwards in time, in 3D

Interactive Figure 1 (Swiggum et al. 2024)

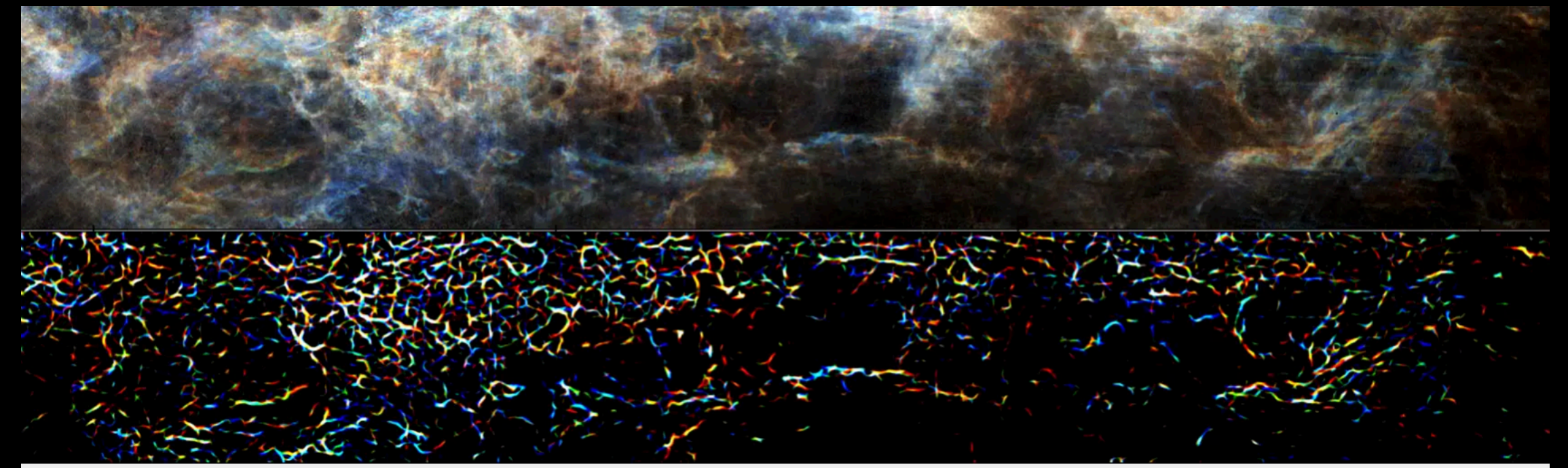
Click to toggle traces on/off

- Parent cluster sample
- Cr135 family
- M6 family
- α Per family
- Sun
- R = 8.12 kpc



Swiggum et al. 2024

The archaeological record of the Milky Way written in neutral atomic hydrogen (HI)



Atomic hydrogen emission from an excerpt from the THOR survey (top) and associated filamentary structures around the... [more]

Image: J. Soler et al. 2020

Soler et al. 2020

Next...

MILKYWAY3D.org



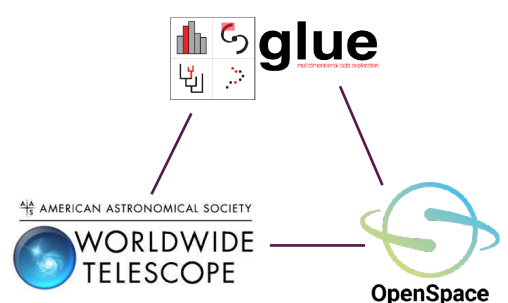
SCAN ME

INFRASTRUCTURE

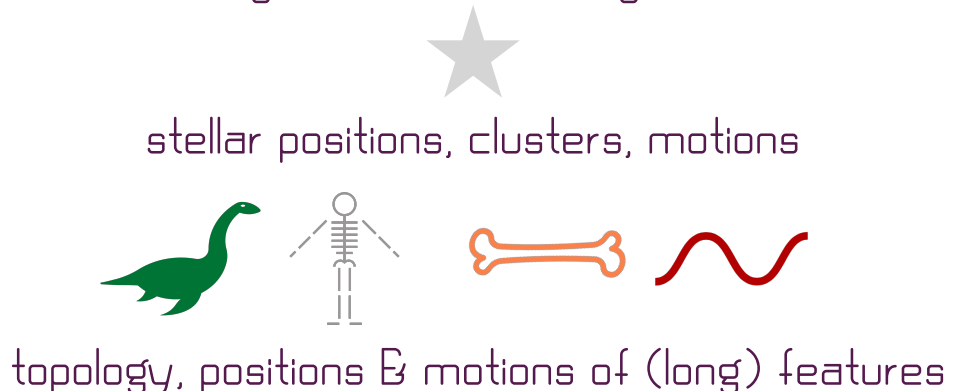
SCIENCE

EDUCATION & OUTREACH

assembling data as a community, using modern, open-source practices



enabling studies of how galaxies turn gas into stars, using...



connecting real research data, software, and science to learners



real-time data exploration anywhere, including in planetaria



"Cosmic Data Stories" teach data science using astronomical data & tools

Lead: Jackie Faherty, AMNH

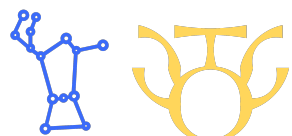
linking position and motion across dimensions, using the plug-in architecture enabled by glue (plug-ins include WorldWide Telescope, OpenSpace & more)



making data accessible online for decades

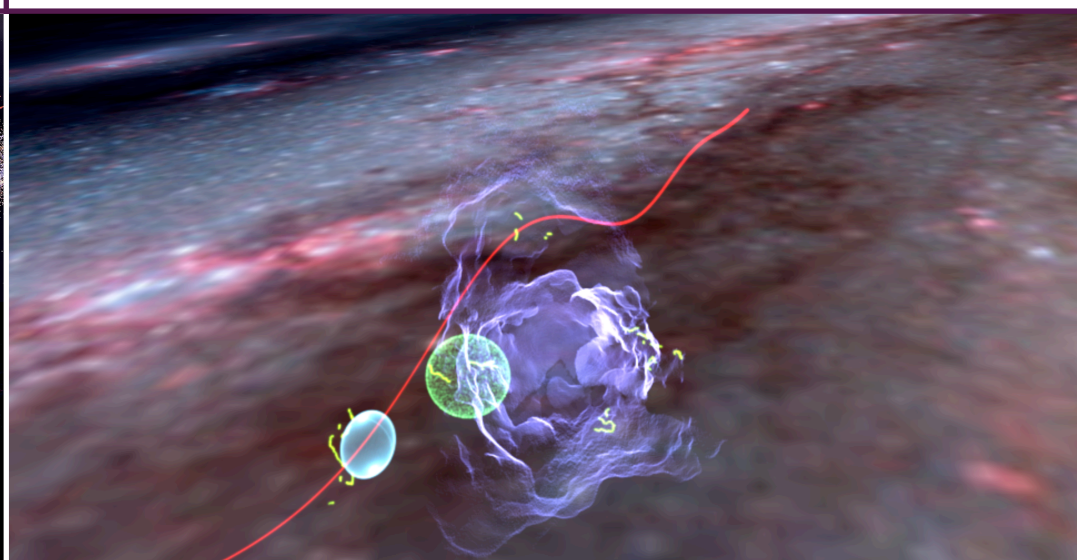
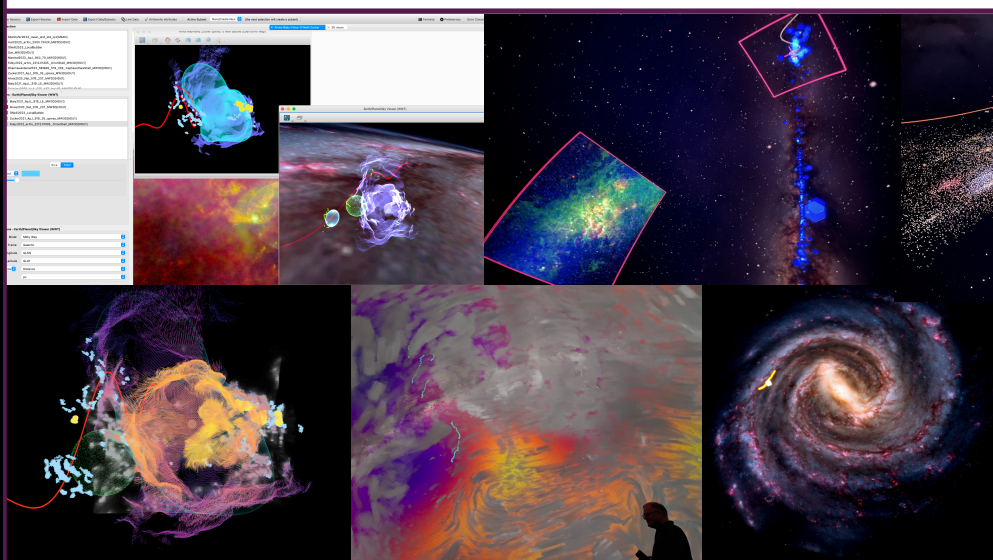
Lead: Alyssa Goodman, CfA

tracers of feedback & magnetic fields



details on star-forming regions...and more!

Lead: Catherine Zucker, CfA



The Sun's Neighborhood in the Milky Way

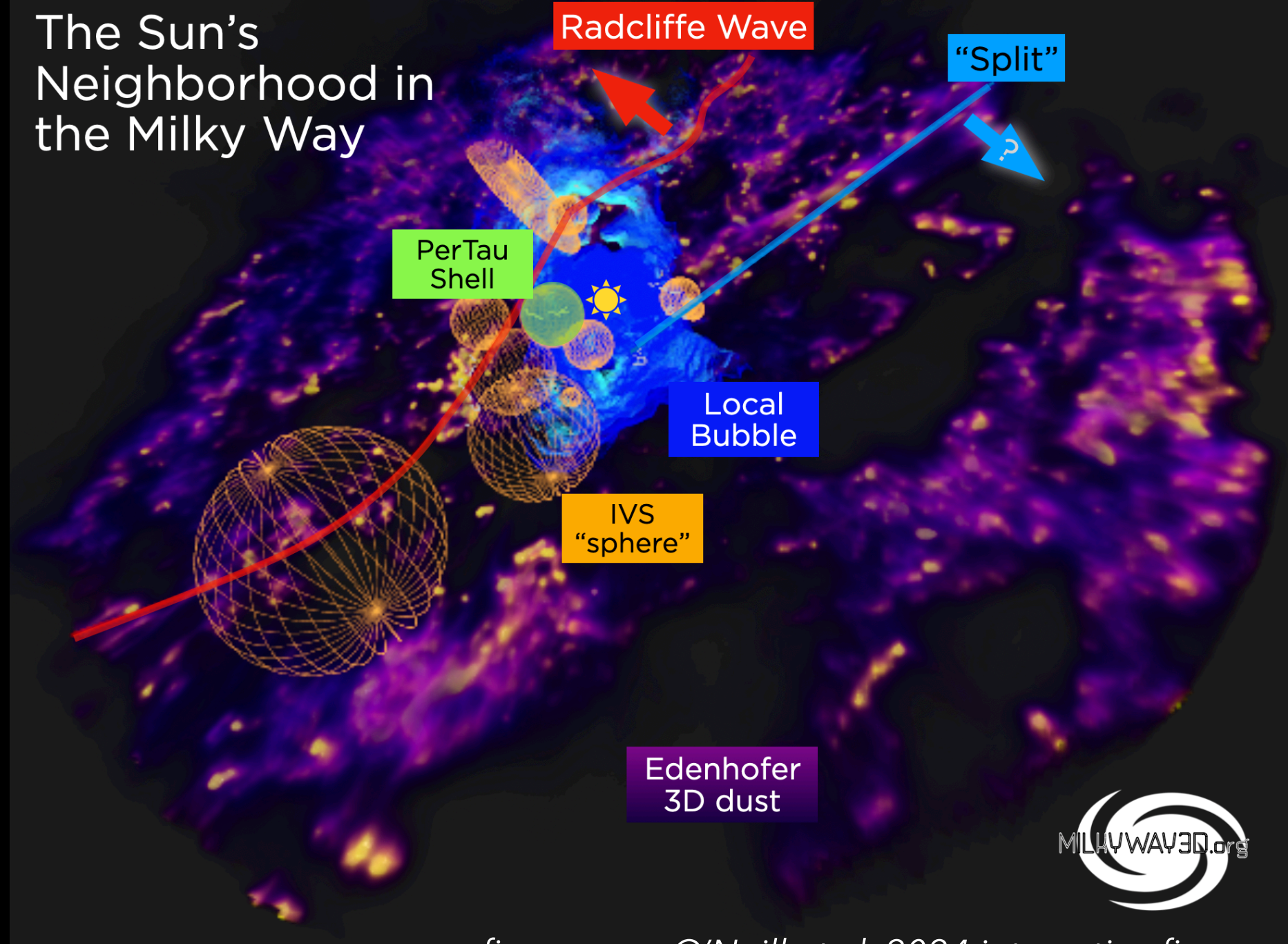
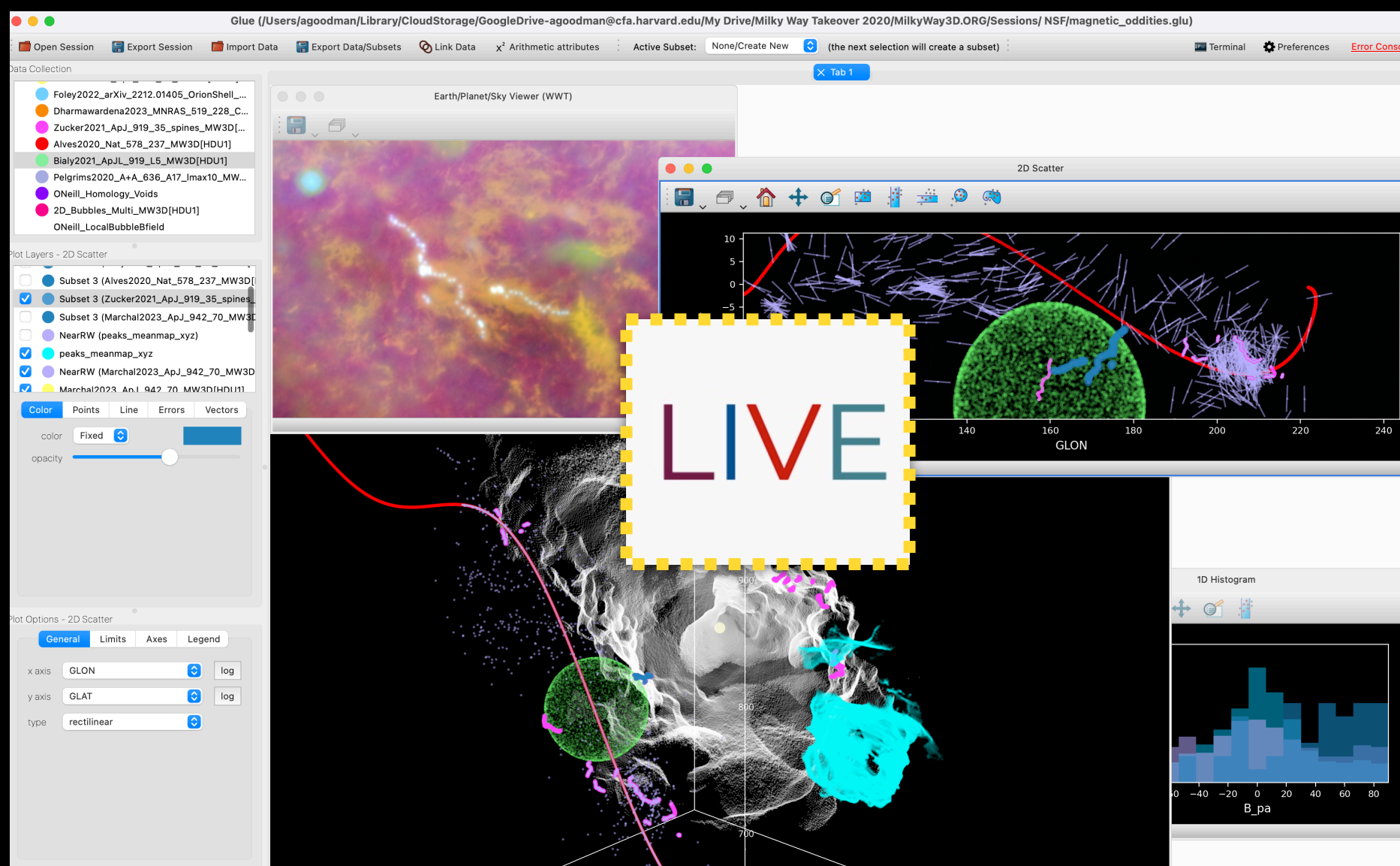


figure uses O'Neill et al. 2024 interactive figure



TEAM: Harvard/Smithsonian CfA (Jonathan Carifio, Alyssa Goodman, Ralf Konietzka, Theo O'Neill, Patricia Udomprasert, Catherine Zucker), AMNH (Brian Abbott, Micah Acinapura, Carter Emmart, Jackie Faherty); Linköping University (Alex Bock); University of Vienna (Joao Alves, Sebastian Ratzenbock); glue solutions, inc./Aperio (Thomas Robitaille); University of Wisconsin, Whitewater (Bob Benjamin), STScI/Johns Hopkins (Josh Peek), Max Planck IfA (Gordian Edenhofer); Northeastern University (Michelle Borkin); and YOU?!



Join us, contribute, and yes, you get a T-Shirt.

Thank you!

tinyurl.com/HHSF-Miro

The Miro board is a collection of astronomical research materials, including:

- NGC6946 (d = 6 Mpc):** Two panels showing the galaxy in different wavelengths.
- Atomic filament orientation:** A diagram showing filament structures in galaxies.
- Marasco et al. 2013:** A diagram illustrating the formation of spiral arms.
- IC342 (HI overlay):** A map showing the distribution of neutral hydrogen in the IC 342 galaxy group.
- CO abundance plot:** A graph showing CO abundance as a function of τ_{CO} and CR ionization rate.
- Other images:** Various spiral galaxies, star-forming regions, and data plots.

The board is annotated with letters and numbers in colored circles:

- Letters:** A, E, H, J, M, N, R, T, Y.
- Numbers:** 1, 2, 3, 6, 10.

Other notable elements include a pink scribble, a blue speech bubble with the name "Noé Brucy", and a large blue circle with the number "10".

NGC6946 (d = 6 Mpc)

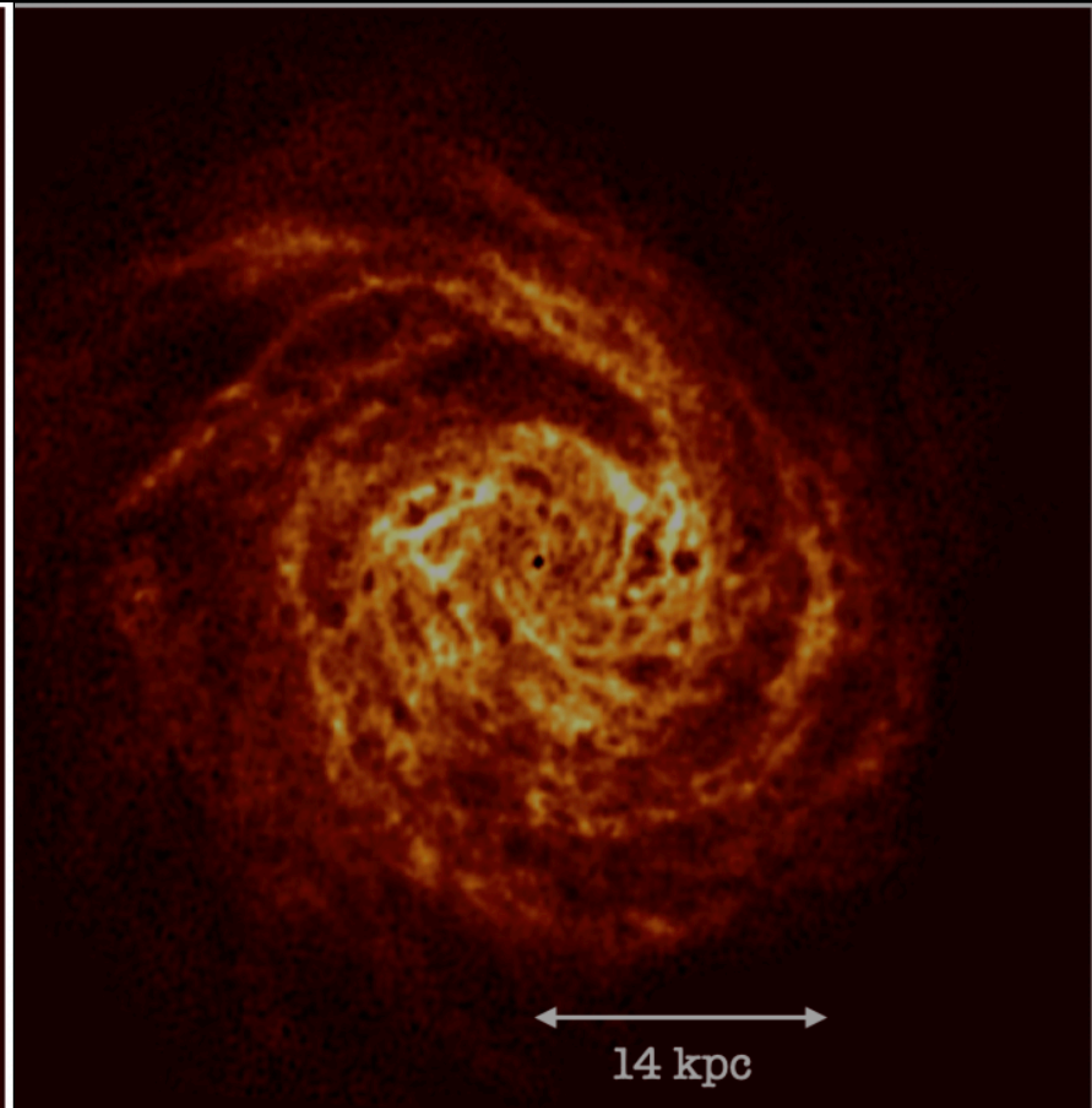
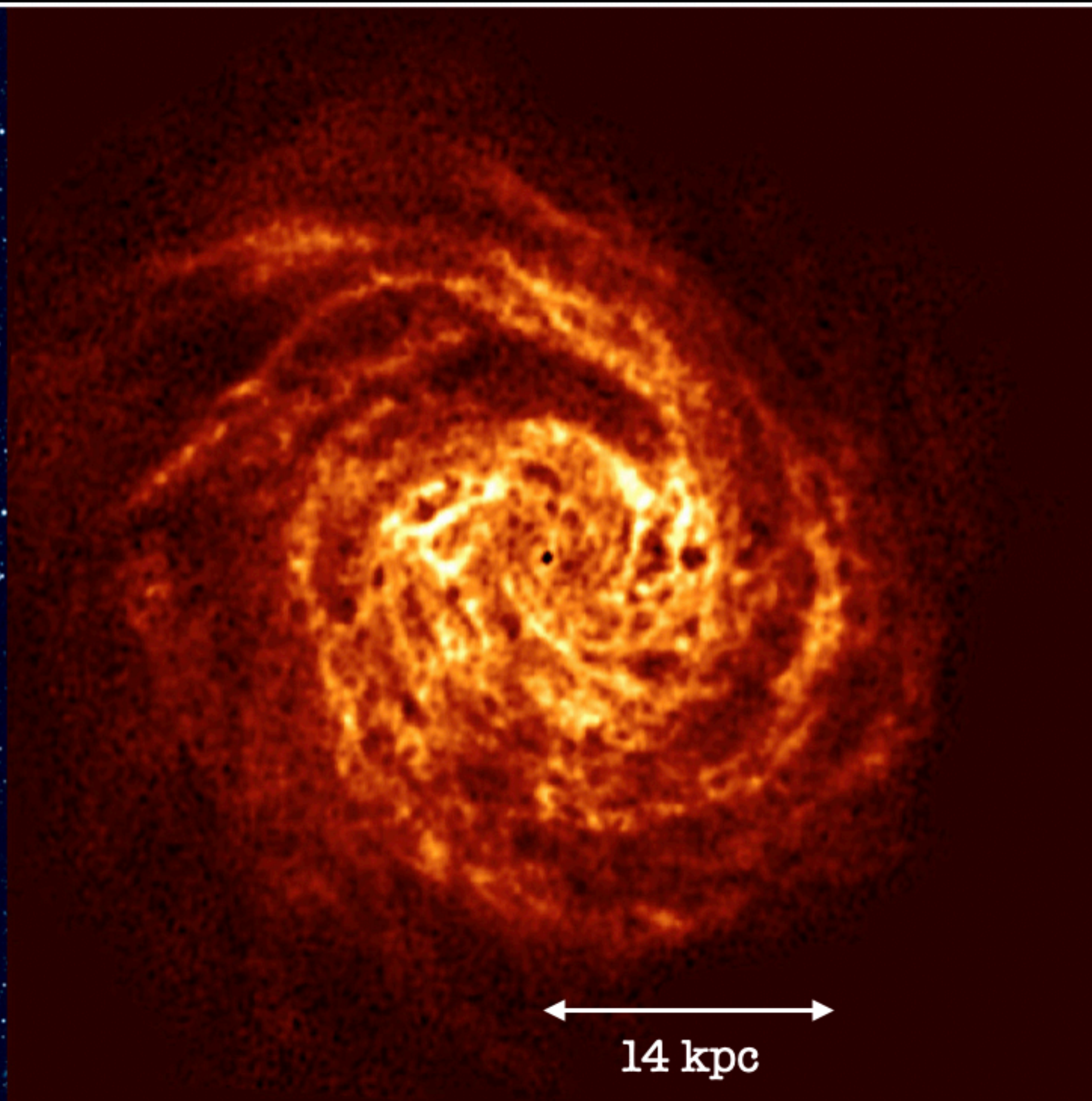
= 6 Mpc)

Thank you!

14 kpc

14 kpc

14 kpc



SOC & LOC

Henrik Beuther (chair!!!)

Catharina Fairchild

Annie Gao

Philipp Girichidis

Christelle Hiemstra

Ralf Klessen

Eric Koch

Shanghuo Li

Theo O'Neill

Simon Reyes

Molly Wells

Daniel Ziegler

Thank you!

and to

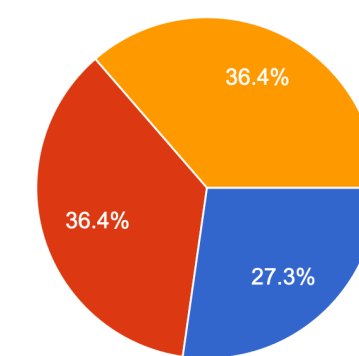
Thomas Henning

who really

“invented” HHSF!

and to **YOU**

How am I helping...
11 responses



- Already gave you an abstract, and that's all I have time for now.
- Already gave you an abstract, and now I'll fill out a 5-minute form (best for posterity!)
- No abstract, but I'll fill out the 5 minute form